IRIS: A stratospheric water isotope ratio spectrometer based on optical-feedback cavity enhanced absorption

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Brief project history: 1999: Dutch NSF grant for whole air, cryogenic sampling project 2001: Proposal redefined: in-situ, spectroscopic detection

Increased temporal and thus spatial coverage

·No need for cryogens

·Reduce chance of sampling related (systematic) errors

2003/1: PhD (RI) started, collaboration with Grenoble initiated

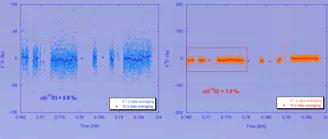
2003/7: Assured access to NASA DC-8 (HJJ)

2004/6: First successful DC-8 test flights at Dryden

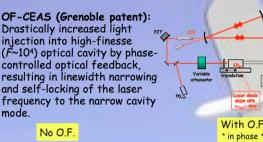
2005/6: Second generation device for stratospheric flights

DC-8 device performance:

Assuming constant water vapor conditions during level flight, precisions are 9% for $\delta^{2}\text{H},$ 3% for $\delta^{17}\text{O},$ and 1% for $\delta^{18}\text{O}$ (10 s averaged data, ~200 ppm).



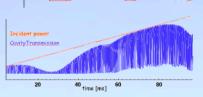
Second generation device: Shot noise limited detectors, better laser (power, tuning), reduced sensitivity to acoustic noise, autonomous operation with PC104 computer, higher flow rate (cell refreshed ~10 s-1), vapor phase only inlet

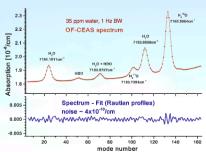




Cavity transmission

The absorption lines are superimposed on the cavity transmissions and appear enhanced as if observed as the output of a multiplepass cell with an effective length given by ~F·L, of the order of 10 km. The spectrum is recorded cavity mode-bycavity mode, separated by the cavity FSR (~150 MHz), such that noise on the

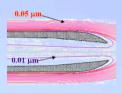




Vapor phase inlet:

Designed for particle-cut-off at 10 nm and no wall contact.

·shaped backward facing inlet ·subsampling near instrument



Mounting Plate



Technical specifications:

Optical head ~70x35x18 cm³, PC ~13x16x21 cm³. Total weight (include. Pump and inlet) ~ 45 kg. Power consumption 175 W typical, 280 W max.

Acknowledgements:

completely negligible.

frequency scale is

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